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Website: <http://www.physik.uzh.ch/lectures/agr/>

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**Exercise 1** [Gravitational field of a moving particle] (5 points)

Consider a particle of mass  $M$  moving with constant velocity  $|\mathbf{V}| \ll c$ . Calculate the gravitomagnetic potential and the equation of motion for a test particle  $m$  in the field generated by  $M$  to order  $\mathcal{O}(v/c)$ . *Hint: Start in the  $M$  particle rest frame  $\Sigma'$  and then perform a coordinate transformation to the global frame  $\Sigma$ . Since we are working at  $\mathcal{O}(h)$  the transformation is nothing than a LORENTZ boost, which can be further approximated since  $V \ll c$ .*

**Exercise 2** [Particles in the field of a gravitational wave] (3 points)

Show that the curves  $r = r(\varphi)$  described by

$$r^2(\varphi) = R^2 \begin{cases} 1 - 2h \cos(2\varphi) \cos(\omega t) \\ 1 - 2h \sin(2\varphi) \cos(\omega t) \end{cases} \quad (1)$$

for  $h \ll 1$  are ellipses. How is the eccentricity  $\epsilon$  related to  $h$ ? *Hint: Start from the expression of an ellipse in polar coordinates and assume  $\epsilon \ll 1$ .*

**Exercise 3** [Gravitational Bremsstrahlung] (6 points)

The gravitational wave analogue of Bremsstrahlung can be generated by a small mass  $m$  scattering off a large mass  $M \gg m$  with impact parameter  $b$ . Assume that the large mass sits at  $(0, 0, 0)$ , that  $E = 0$  (parabolic orbit) and that the orbit lies in the x-y-plane. The parabolic trajectory of the particle  $m$  in polar coordinate is described by:

$$r(\varphi) = \frac{2b}{1 + \cos(\varphi)} \quad \dot{\varphi} = \sqrt{\frac{M}{8b^3}} [1 + \cos(\varphi)]^2. \quad (2)$$

Calculate the gravitational wave amplitude  $h_{\mu\nu}$  starting from<sup>1</sup>:

$$\gamma_{ij}(t, \mathbf{x}) = \frac{2G}{r} \frac{d^2 I_{ij}(\varphi(t_r))}{dt^2}, \quad (3)$$

where  $\gamma_{\mu\nu} = h_{\mu\nu} - 1/2 h \eta_{\mu\nu}$ ,  $t_r$  is the retarded time and  $I_{ij}(t) = \int T_{00}(t) y^i y^j dy^3$  is the quadrupole tensor. *Hint: It is not necessary/required to express the final  $h_{\mu\nu}$  as a function of time, it is enough to show the dependence on  $\varphi$ , but use the expression for  $\dot{\varphi}$  to simplify the evaluation of the time derivative in equation (3).*

<sup>1</sup>See Carroll for derivation and restrictions of this formula, <http://arxiv.org/abs/gr-qc/9712019v1>.