Maintenance of critical industrial equipment using cosmic muon radiation.

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Outline of the talk

→ Cosmic muons: numbers, orders of magnitude and detectors.

→ Muography: Principles, types, applications.

Maintenance of industrial equipment using muography: Muon Systems.

→ The future of Muography and next steps.

Cosmic rays: composition, flux, energy spectrum.



Cosmic rays: composition, flux, energy spectrum.

- Constant flux of high energy particles bombarding the Earth from the space.
- Composed mainly by protons (98%), α particles (1.8%) and others (< 0.2%)
- → Cosmic rays are the most energetic particle ever seen so far.



The Oh-My-God Particle 3x10⁸ TeV (1991, several since then)

http://www.cosmic-ray.org/reading/flyseye.html#SEC10

Cosmic muons: flux and energy spectrum.

- Muons are generated mostly from pion decays in the atmosphere.
- The flux of muons is mostly proportional to $\cos^2(\theta)$ (angle with the vertical).
- Quickly falling spectra \rightarrow average of **3 GeV** when integrating solid angle.



Rule of thumb at the surface:





Muon interaction with matter: ionization.

- Ionization is one of the most frequent processes for cosmic muons.
- → Energy loss depends on the Z, density, and size of the crossed object.
- → The Range is the distance for which the particle looses all the energy.



depth (km.w.e.)

Muons interaction with matter: multiple scattering.

- Coulomb scattering deviates the direction of muons when crossing matter.
- → Scattering angle depends also on the Z, density, and size of the material.
- → Angular distribution is approximately gaussian (with some tails).



Muon Energy/GeV	Material	Width/cm	Angle/mrad
1	Iron	1	10
10	Iron	1	1
1	Iron	10	34
10	Iron	10	3
1	Lead	1	19
10	Lead	1	2
1	Lead	10	64
10	Lead	10	6
1	Uranium	1	26
10	Uranium	1	3
1	Uranium	10	88
10	Uranium	10	9

Muon detectors using different technologies.

- Multiwire proportional chambers and drift tubes (gaseous detectors).
- Scintillitator detectors (of different kinds) + photomultipliers.
- → Sillicon-based detectors: pixel-like, strip-like, etc.
- Important parameters: spatial resolution (granularity), efficiency, acceptance.



http://cds.cern.ch/record/942973 -

CIEMAT assembly of DT chambers

Muography: principles, types, and applications.



An old (documented) Muography application.

- Muography uses cosmic muons to obtain images of unaccesible places.
- First attempt by Luis Álvarez in 1970 to find hidden chambers in pyramids.

Fikhry Hassan, Dennis Iverson, Gerald Lynch, Zenab Miligy

The three pyramids of Giza are situated a few miles southwest of Cairo, Egypt. The two largest pyramids stand within a few hundred meters of each other. They were originally of almost exactly the same height (145 meters), but the Great Pyramid of Cheops has a slightly larger square base (230 meters on a side) than the Second Pyramid of Chephren (215.5 meters on a side). A photograph of the pyramids at Giza is shown as Fig. 1. Figure 2 shows the elevation cross sections of the two pyramids and indicates the contrast in architectural design. The simplicity of Chephren's pyramid, compared with structure of his father's the elshorate Great Pyramid, is explained by archeologists in terms of a "period of ex-perimentation," ending with the construction of Cheops's pyramid (1). (The complexity of the internal architecture of the pyramids increased during the Fourth Dynasty until the time of Cheops and then gave way to quite

simple designs after his time.) An alternative explanation for the sudden decrease in internal complexity from the Great Pyramid to the Second Pyramid suggested itself to us: perhaps Chephren's architects had been more successful in hiding their upper chambers than were Cheops's. The interior of the Great Pyramid was reached by the tunneling laborers of Caliph Ma-

g invertee are affiliated with the Joint Pyra-the United Arab Republic and the of America. They reside either in Arab Republic, or in Berkeley, meter is adapted from an adangular resolution had been used in

Search for Hidden Chambers in the Pyramids

The structure of the Second Pyramid of Giza is determined by cosmic-ray absorption.

Luis W. Alvarez, Jared A. Anderson, F. El Bedwei, James Burkhard, Ahmed Fakhry, Adib Girgis, Amr Goneid, Ali Hilmy Moussa, Mohammed-Sharkawi, Lauren Yazolino

> mun in the 9th century A.D., almost largest pyramids. Since that time we 3400 years after its construction. Of have accumulated accurate angular our group only Ahmed Fakhry (author measurements on more than a million of The Pyramids, professor emeritus of cosmic-ray muons that have penetrated archeology, University of Cairo, and an average of about 100 meters of member of the Supreme Council of Archeology, Cairo) was trained in arin the Belzoni Chamber, cheology. As laymen, we thought it not unlikely that unknown chambers might still be present in the limestone above Proof of the Method the "Belzoni Chamber," which is near the center of the base of Chephren's and Pyramid, and that these cham in an exploratory mode, it is essential bers had survived undetected for 4500 that the capabilities of the technique years. [We learned later that such ideas be demonstrated on a known system. had occurred to early 19th-century in-We gave serious consideration to a vestigators (2), who blasted holes in the proposal that the cosmic-ray detectors be tested first in the Queen's Chamber pyramids with gunpowder in attempts of the Great Pyramid, to demonstrat

to locate new chambers.] In 1965 a proposal to probe the Second Pyramid with cosmic rays (3) was sent to a representative group of cosmic-ray physicists and archeologists with a request for comments concern ing its technical feasibility and archeological interest. The principal novelty of the proposed cosmic-ray detectors

involved their ability to measure the angles of arrival of penetrating cosmicupper chamber would give a large signal. It was apparent that the only ray muons with great precision, over a large sensitive area. The properties of untested feature of the new technique involved the magnitude of the scatte the penetrating cosmic rays have been ing of high energy muons in solid mat sufficiently well known for 30 years to suggest their use in a pyramid-probing ter. (An anomalously large scattering would nullify the high angular resolu experiment, but it was not until the invention of spark chambers with digition that had been built into the de tectors, in the same way that frosted tal read-out features (4) that such a glass destroys our ability to see distant use could be considered as a real pos-sibility. [Cosmic-ray detectors with low objects.) We had no reason to doubt the calculated scattering, but we were

anxious to be able to demonstrate to 1955 to give an independent measure our colleagues in the U.A.R. Depart-SCIENCE, VOL. 167

of the thickness of rock overlying ar underground powerhouse in Australia's Snowy Mountains Scheme (5)].

The favorable response to the pro-posal led to the establishment by the

United Arab Republic and the United States of America of the Joint U.A.R.-

U.S.A. Pyramid Project on 14 June

1966. Cosmic-ray detectors were in-

stalled in the Belzoni Chamber of the Second Pyramid at Giza in the spring

of 1967 by physicists from the Ein Shams University and the University

of California, in cooperation with an

cheologists from the U.A.R. Depart-

ment of Antiquities. Initial operation

had been scheduled for the middle of June 1967, but for reasons beyond our

control the schedule was delayed for

several months. In early 1968 cosmic

ray data began to be recorded on mag

setic tape in our laboratory building

a few hundred meters from the two

imestone on their way to the detectors

Before any new technique is used

that the King's Chamber and the Grand

Gallery could be detected. But this suggestion was abandoned because the

King's Chamber is so close to the

Queen's Chamber and because it sub

tends such a large solid angle that ear

ments had already shown that the

on) cosmic-ray experi

nalously large scattering

lier (low resolution

Fig. 1 (top right). The pyramids at Giza. From left to right, the Third Pyramid of Mycerinus, the Second Pyramid of Che-phren, the Great Pyramid of Cheops. [© National Geographic Society]

ment of Antiquities in a convin manner that the technique really worked as we had calculated. For this purpose we required as our test objects not large features that were nearby but, instead, small features separated from the detectors by the greatest possible thickness of limestone. Fortunately, such features are available in the Second Pyramid: the four diagonal ridges that mark the intersections of neighbor-ing plane faces were farther from the detectors than any other points on the individual faces. (From now on, we will refer to these ridges as the "cor

From the known geometry of the Second Pyramid, the trajectories of cosmic-ray muons that pass through a point on a face 10 meters from a corner and then down to the detectors can be shown to traverse 2.3 fewer meters of limestone than do muons that strike the corner. They should therefore arrive with 5 percent greater intensity than the muons from the corner. Such an increase in intensity, corresponding to such a decrease in path through the mestone, is about half of what would be expected to result from the presence of a chamber of "typical size" (5 me-ters high) in the pyramid. Since such a chamber would necessarily be closer to the detectors, it would for these two reasons be a much "easier object to see" than the corner.

The detection equipment was there fore installed in the southeast corner of the Belzoni Chamber, with the expectation that it would first show the corners in a convincing manner, so that the presence or absence of unknown chambers could later be demonstrated to the satisfaction of all concerned. In September 1968 the IBM-1130 computer at the Ein Shams University Computing Center produced the data

Fig. 2 (bottom right). Cross sections of (a) the Great Pyramid of Cheops and (b) the Granbers: (d) Smooth Insteamed cap, (d) the Betoni Chamber, (C) Betoni's en-trance, (D) Howard/Sysie startance, (E) descenting paragersay, (E) ascending (H) Grand Gallery, (D) King's Chamber, (H) Gallery, (H) King's Chamber, (H) Gallery, (H)





"We have no doubt that we could detect a King's Chamber anywhere above the Belzoni Chamber within a cone of half-angle 35 degrees from the vertical.

If the Second Pyramid architects had placed a Grand Gallery, King's Chamber, and Queen's Chamber in the same location as they did in Cheops's Pyramid, the signals from each of these three cavities would have been enormous.

We therefore conclude that no chambers of the size seen in the four large pyramids of the Fourth Dynasty are in our "field of view" above the Belzoni Chamber."

http://lappweb.in2p3.fr/~chefdevi/Work_LAPP/Arche/alvarez_70.pdf

Two categories on muography applications.

Two categories according to the physics process: absorption vs. scattering.

Absorption Muography

- Muon flux measured as a function of the direction.
- Differential transmittance "T".
- Need Large Masses + Long exposure times.
- → Applications: vulcanology, geology, archeology...
- Only one detector needed.



Scattering Muography

- Scattering position and angular shift.
- Smaller Masses + shorter exposure times.
- → Applications: security, industry, etc.
- At least two detectors are needed.



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The inverse problem in muon tomography.

→ Tomography is a classic "inverse problem" type of challenge.



- $\textbf{\textbf{+}}$ Where the physical process depends on some parameters λi
- → And where we know how to solve the forward problem:
 - Given "x" and " λ_i " we know how to estimate "y".
- The inverse problem consists of estimating " λ_i " when having "x" and "y"
- → BUT: in this case f() is not a deterministic function \rightarrow it is a probability density.

The simple geometric approach: POCA

- This complex problem can be approximated doing some physical assumptions
 - The scattering occurs only at one point during the whole trajectory.
 - Works very well for scenarios that are essentially empty with some high density chunks.
- → The scattering is assumed to be produced in the Point of Closest Approach.
- And the density is assumed to be in the places where there is scattering.

POCA point

Point of Closest Approach (POCA)

The "I wish" Maximum Likelihood solution

- → Since the forward problem is known one could thing about applying Wick's theorem
 - Fill the target volume with a set of voxels of random density.
 - → For every input muon calculate the pdf of the output muon using simulation.
 - Evaluate the output muon in the pdf and accumulate in the likelihood function.
 - Minimize the likelihood to get the best densities.
- This is just impossible due to time/computing resources.



 $L(x, y; \lambda) = \prod pdf(y_i; x_i, \lambda)$ $minL(x, y; \lambda)$

The "poor-man" Maximum Likelihood solution

- The complexity of the previous procedure reduces drastically if:
 - The propagation of the muon is approximated as a product of gaussians in the voxels.
 - The voxels crossed for each muon are fixed and approximated (under some assumption).
- Under this conditions an analytical likelihood can be obtained:

$$P(H_{ij} \mid \lambda) = \frac{1}{2\pi |\Sigma_{ij}|^{1/2}} \exp\left(-\frac{1}{2} H_{ij}^T \Sigma_{ij}^{-1} H_{ij}\right) \qquad \log(P(H \mid \lambda)) = \sum_{j \le N} \sum_{i: L_{ij} \ne 0} \left(-\log \lambda_j - \frac{H_{ij}^T A_{ij}^{-1} H_{ij}}{2\lambda_j p_{r,i}^2}\right) + C \quad (34)$$

- However the procedure is also quite time expensive and
- → The results are not dramatically better than POCA:
 - Since in the end a POCA assumption is made on the trajectory of the muons.

Application to vulcanology



Monitoring of Vesuvious vulcano.



X_1	125	0
<i>X</i> ₂	235	+110
<i>X</i> ₃	360	+125
X_4	385	+25
X_5	985	+600
<i>X</i> ₆	1010	+25
<i>X</i> ₇	1135	+125
X_8	1245	+110
X_9	1370	+125
<i>X</i> ₁₀	1495	+125
<i>X</i> ₁₁	1605	+110
<i>X</i> ₁₂	1730	+125
<i>X</i> ₁₃	1855	+125
<i>X</i> ₁₄	1965	+110

Only 1 scintillator detector.

Exposure time ~ year.

Dinamic studies on magma convenction.





https://doi.org/10.1098/rsta.2018.0142

Monitoring of Asama & Iwujima vulcanos.

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Application to mineral mines.

https://doi.org/10.1098/rsta.2018.0061

Search for uranium/lead clusters in McArthur River (Canada) and Pend Oreille (USA)

Canadian Private company:

CRM Geotomography Technologies, Inc.,

Grid of scintillator detectors in the galleries.



Underground mapping for pipe location.

https://doi.org/10.1098/rsta.2018.0133

Israel Private company:

Lingacom S.L,

Cylindrical-scintillator detectors.





Application to archeology/civil engineering.

Search for hidden galleries in Mount Echia (Bourbon Tunnel)



https://doi.org/10.1098/rsta.2018.0057

Only 1 scintillator detector.

Exposure time ~ 26 days.





computer and ancillary

Two drift tube chambers located outside and inside the dome.

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Application to border security.

Truck inspection using cosmic muons.

Using CMS-like chambers at Torino.

For this application: 4 chambers.

Detection of a 7cmx7cmx7cm lead box.



Decision Science

First commercial system.

First contract in Singapur to install portal.

Multi-technology detection portal.

https://decisionsciences.com/

Application to the nuclear industry.





Nuclear waste characterization.

Lynkeos Technology Ltd (University of Glasgow)

UK Nuclear Decommissioning Authority.

Commercial Scintillator detector.

https://doi.org/10.1098/rsta.2018.0048

Fuel cask monitoring.







https://doi.org/10.1098/rsta.2018.0052

Exposure times of about 7 hours.

Effective detection of missing fuel.

Maintenance of industrial equipment using Muography: Muon Systems



Degradation of the interior of industrial equipment.

Impact in the operation due to the changes in the width in critical parts of the installation: pipes, cauldrons, blast furnaces.



Humidity, corrosion, high pressure and temperature.



Preventive stops (blast furnaces). Measurement of the width.



Changes in the width of the walls. Alteration of mechanical conditions.



High economical and potentially energetic cost of unexpected stops.

Muon Systems & Instituto de Física de Cantabria

Muon Systems: a young Muography company.

- → Founded in 2017 by 2 members using private capital.
- → Based on Bilbao (Edificio Ilgner, Altos Hornos de Vizcaya).
- Currently 3 employees + 1 consultant.



https://muon.systems/



New input in preventive maintenance.

- → Using multiple scattering to reconstruct the inner geometry of the equipment.
 - High penetration power.
 - → No physical contact with the facility.
 - No need of stopping the productive process.
 - Possibility to perform continuous monitoring for critical cases.



Financial/technical support from many institutions.



Area 1: Innovative projects.

ELISKO JAURLARITZA GOBIERNO VASCO PEDROMANYO GAMPON EN LEMPONYO DE CEMANOLIO ECONOMICO Y COMPACTIVICADO

Ekintzaile: Innovative projects.



Best national project in manufacturing and materials CDTI – NEOTEC 2017



Progress in control of processes and maintenance.



Accelerated by FDE.



Sidenor challenge.

A lot of interest from many different companies.

- Different problems ranging from huge facilities to small equipment.
- → In this presentation I will be focusing on the problem of insulated pipes.



Measurement of the width of an insulated pipe.



DOI:10.1098/rsta.2018.0054

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Measurement of the width of an insulated pipe.

- Measure multiple scattering between the upper and lower trajectories.
- → Point of Closest Approach → most likely scattering center.



https://nuclear.llnl.gov/simulation/doc_cry_v1.7/cry.pdf

Reconstructed images for different widths.

Muon Systems Position of POCA hits [m] Z 20 700 600 10 500 400 0 300 -10 200 -20 100 20 Y [cm] -20 -10 0 10

Thickness = 1.2 cm

Thickness = 1.6 cm



Thickness = 1.8 cm



Thickness = 2.6 cm



Thickness = 2.0 cm



Thickness = 2.4 cm



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Muon

ystems

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HISICa

de

Cantabria

- → The loss of resolution makes it impossible to see the insulation layer.
- The steel wall is dense enough to be detected without problems.



Perfect resolution

4 mm resolution

Muon Systems



Thickness = 2.0 cm



Thickness = 1.6 cm



Thickness = 2.4 cm



Thickness = 1.8 cm



Thickness = 2.6 cm



Maintenance of critical industrial equipment using cosmic muon radiation.

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- The loss of resolution makes it impossible to see the insulation layer.
- → The differences in the radial distribution for the steel are still significant.
- → Algorithm (*) looking for statistical compatibility between target and template.
- → Reach resolutions of about 1-2 mm with 1 hour measurement.



Accuracy

(*) Based on **Deep Learning Convolutional Network** no restricted to azymuthal asymmetry.

First prototype of (small) muon chambers.



First prototype of (small) muon chambers.







Maintenance of critical industrial equipment using cosmic muon radiation.

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Commissioning of chambers (I).

Light DAQ and DQM running on Raspberry-Pi and sending data through wireless.

Good general functioning of individual chambers for wires.

Still low detection efficiency (~60%) due to multiplexing and strips.



Commissioning of chambers (II).

First tracks reconstructed successfully with a high trigger efficiency (>95%)

Low detection efficiency in the strips coordinate (accumulated from individual chamber).



Detection of a 16 mm lead layer.

- Good agreement between data and simulation.
- Also good discrimination between the two scenarios.



Lessons learned from 1st prototype.

- Good functioning and response of the wires \rightarrow high efficiency system.
- Strips discarded because of multiple strip activation per event.
- → Good functioning of electronics and trigger system (very high efficiency).
- → Need to scale the system to increase the acceptance.



The future of muography and next steps.



Scaling detectors to 1 squared meter dimension.





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Improving reconstruction algorithms (I)

- Consider the quartiles of distributions of the following parameters as input to DNN
 - Entrance position and direction of the muon. (x, y)
 - Deviation in position and direction as measured by second muon. (x,y)
 - Products of these variables.
- → Consider the density of every voxel as the output of the DNN.



Improving reconstruction algorithms (II)

Map positions and directions in the detectors to integers.



- → In this way all possible tracks are indexed and the input and output are encoded.
- → The number of events found on each index is fed as the input of the DNN.
- The output is the density per voxel in the output geometry.







Next steps for Muon Systems.

- Detector construction will be ready in the next months.
- Expect to perform commissioning in the lab before the Summer.
- → Plan to do an on-site pilot test in Dynasol in Santander to measure Pipes.



International consortium of Muographers.

- Muography is slowly entering into a commercial phase.
- → Attempts hosted by the IAEA and the Royal Society to create a consortium.
 - Idea is to work together to promote the technology worldwide.



THANK YOU



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