



Two-loop Yukawa corrections to double Higgs production

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in collaboration with Joshua Davies, Go Mishima, Matthias Steinhauser and Hantian Zhang | Oktober 18, 2022



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Higgs Self Coupling



Standard Model Higgs potential:

$$V(H)=rac{1}{2}m_H^2H^2+\lambda extsf{v}H^3+rac{\lambda}{4}H^4$$
 , where $\lambda=m_H^2/(2 extsf{v}^2)pprox 0.13$

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- Want to measure λ , to determine if V(H) is consistent with nature.
 - Challenging! Cross-section $\approx 10^{-3} \times H$ prod.
 - $-3.3 < \lambda/\lambda_{SM} < 8.5$ [CMS '21]
- λ appears in various production channels:



Gluon fusion – dominant, 10x

VBF

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tt associated production

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Gluon Fusion

Leading order (1 loop) partonic amplitude:



- \mathcal{F}_{tri} contains the dependence on λ at LO
- Form factors:
 - LO: known exactly
 - Beyond LO... no fully-exact (analytic) results to date
 - QCD: numerical evaluation, expansion in various kinematic limits
 - EW: first steps: this work (HE)
 - (see also HTL considerations)

[Glover, van der Bij '88]

[Davies, Mishima, Schönwald, Steinhauser, Zhang '22] [Mühlleitner,Schlenk,Spira '22]

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gg ightarrow HH Beyond LO

NLO QCD:

large-mt

numeric

- large-m_t + threshold exp. Padé
- high-energy expansion
- small- p_T expansion

NNLO QCD:

- large-m_t virtuals
- HTL+numeric real ("FTapprox")
- large-m_t reals

N3LO QCD:

- Wilson coefficient C_{HH}
- HTL

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[Spira '16: Gerlach, Herren, Steinhauser '18]

 $\underset{\circ}{\text{Conclusion and Outlook}}$

[Chen, Li, Shao, Wang '19]

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 [Dawson, Dittmaier, Spira '98] [Grigo, Hoff, Melnikov, Steinhauser '13]
 [Borowka, Greiner, Heinrich, Jones, Kerner, Schlenk, Schubert, Zirke '16] [Baglio, Campanario, Glaus, Mühlleitner, Spira, Streicher '19] [Gröber, Maier, Rauh '17]
 [Davies, Mishima, Steinhauser, Wellmann '18,'19] [Bonciani, Degrassi, Giardino, Gröber '18]

[de Florian, Mazzitelli '13] [Grigo, Hoff, Steinhauser '15, Davies; Steinhauser '19] [Grazzini, Heinrich, Jones, Kallweit, Kerner, Lindert, Mazzitelli '18] [Davies, Herren, Mishima, Steinhauser '19 '21]



gg ightarrow HH Beyond LO



Total cross section (14TeV):

	σ_{LO}	$\sigma_{\it NLO}$	σ_{NNLO}
B-i HTL	-	$38.32^{+18.1\%}_{-14.9\%}$	$39.58^{+1.4\%}_{-4.7\%}$
FTapprox	_	34.25 ^{+14.7%}	36.69 ^{+2.1%}
Full	19.85 ^{+27.6%} _20.5%	32.88 ^{+13.5%}	-

[Borowka, Greiner, Heinrich, Jones, Kerner '16]

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Electroweak Corrections



As we investigate NNLO QCD and beyond, we should consider NLO EW:

$$\mathcal{M} \sim \alpha_{s} \alpha_{t} \Big(\mathbf{A}_{1} + \alpha_{s} \mathbf{A}_{2} + \alpha_{t} \mathbf{A}_{3} + \alpha_{t,\lambda,gauge} \mathbf{A}_{4} + \mathcal{O}(\alpha_{s}^{2}, \alpha_{t}^{2}, \ldots) \Big)$$



There are more scales to deal with, compared to the QCD contribution,

- start with $\alpha_s \alpha_t^2$ diagrams with internally propagating Higgs:
 - expansion parameter not small $\alpha_t = \alpha m_t^2 / (2s_W^2 m_W^2) \sim \alpha_s / 2$
 - only planar integrals in this subset

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High-Energy Expansion

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High-Energy Expansion

• The full diagrams depend on a lot of variables:

- *ϵ*, *s*, *t*, *m*_{*t*}, *m*_{*h*}
- complete analytic solution is out of reach
- First, expand around $m_H^{ext} = 0$ (as for QCD):
 - expand amplitude integrals with LiteRed [Lee '14]
- Unlike for QCD the scale " m_H^{int} " remains, from the propagator:
 - complicates the IBP reduction
 - Master Integrals with this many scales are difficult.
- We expand in this scale also, and propose two ways to do it:
 - A: $s, |t| \gg m_t^2 \gg m_H^{int^2} \sim m_H^{ext^2}$,
 - B: $s, |t| \gg m_t^2 \sim m_H^{int^2} \gg m_H^{ext^2}$.





Asymptotic Expansions

Expansion by Regions [Beneke, Smirnov '98] :

- Assign a hierarchy to the dimensionful parameters.
- Reveal all relevant scalings of the integration variables.
- Expand the integrand according to the scalings for each region.
- Integrate the expanded regions.
- Sum the contributions from all regions.



Asymptotic Expansions

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Here:

$$m_t^2 \ll s, |t|:$$
 $m_t^2 \sim \chi m_t^2$



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Asymptotic Expansions

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Here:

$$m_t^2 \ll s, |t|: \qquad m_t^2 \sim \chi m_t^2$$

- Revealing all relevant regions can be a hard task.
- Automatized in the Mathematica package Asy.m [Pak, Smirnov '11].
- Algorithm is based on the α -parameter representation of Feynman diagrams.

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High-Energy Expansion "A"

Option A: asymptotic expansion around $m_H^{int} = 0$:



The two-loop subgraph is a Taylor expansion of the Higgs propagator:

- results in integrals with a massless internal line, scales s, t, m_t .
- IBP reduce with FIRE and Kira
- these coincide with the QCD Master Integrals reuse the old results

[Smirnov '15] [Klappert,Lange,Maierhöfer,Usovitsch '21]

[Davies,Mishima,Steinhauser,Wellmann '18,'19]





High-Energy Expansion "A"

Option A: asymptotic expansion around $m_H^{int} = 0$:



High-Energy Expansion "B"

Option B: expand around $m_H^{int} \approx m_t$,

- simple Taylor expansion, exp not necessary
 - much easier to implement
- IBP reduce resulting integrals, FIRE+Kira

Write Higgs propagator as: $\frac{1}{p^2 - m_H^2} = \frac{1}{p^2 - m_l^2(1 - [2 - \delta]\delta)}$

• expand around $\delta \rightarrow 0$ where $\delta = 1 - m_H/m_t \approx 0.28$.





High-Energy Expansion "B"

Option B: expand around $m_H^{int} \approx m_t$,

- simple Taylor expansion, exp not necessary
 - much easier to implement
- IBP reduce resulting integrals, FIRE+Kira

Write Higgs propagator as: $\frac{1}{\rho^2 - m_H^2} = \frac{1}{\rho^2 - m_t^2(1 - [2 - \delta]\delta)}$

• expand around $\delta \rightarrow 0$ where $\delta = 1 - m_H/m_t \approx 0.28$.

This yields new integral families compared to the QCD computation:

Calculation of the Master Integrals

Master Integrals Results

• all lines have the mass m_t ,

High-Energy Expansion

compute the MIs in the high-energy limit.

We expand to $(m_H^{ext})^4$ and δ^3 .

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Calculation of the Master Integrals

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Equal Mass Limit

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Calculation of the Master Integrals Equal Mass Limit





- The integral families can be obtained by crossings from the graphs shown above.
- We reduce the scalar integrals with Fire [Smirnov '15] and find 140 master integrals. We make sure to reduce to a minimal set by:
 - We apply FindRules on all scalar integrals and run a second reduction.
 - Equating results of both reduction runs reveals non-trivial relations between master integrals of different families.
 - We run a search for master integrals with Kira [Klappert, Lange, Maierhöfer, Usovitsch '21] .
- We make sure to have a 'good' basis with ImproveMasters [Smirnov '20], i.e.:
 - The denominators factor in $\epsilon = (4 d)/2$ and the kinematics.
 - We get rid of spurious poles in *ϵ*, so that we have to calculate only to *O*(*ϵ*⁰).

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How to solve the master integrals?

- Full solution of the master integrals is still very complicated:
 - Solutions depend on 3 scales: s, t, m_t.
 - The master integrals have up to 7 massive internal lines.
 - The solutions have two thresholds at $\sqrt{s} = 2m_t$ and $\sqrt{s} = 3m_t$.
- However: Analytic solutions possible in the high energy region $m_t^2 \ll s, |t|$.

In the following:

- How to obtain a deep expansion utilizing the differential equations?
- How to obtain boundary conditions to solve the differential equations?
- How well does the approximation work?



Deep Expansion



• Establish a system of differential equations for the master integrals in the variable m_t .

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Deep Expansion



- Establish a system of differential equations for the master integrals in the variable m_t .
- Compute an expansion around $m_t = 0$ by:

• Establish a system of differential equations for the master integrals in the variable m_t .

Calculation of the Master Integrals

Deep Expansion

Introduction

- Compute an expansion around $m_t = 0$ by:
 - Inserting an ansatz for the master integrals into the differential equation.

Calculation of the Master Integrals

$$M_n(\epsilon, m_t \rightarrow 0) = \sum_{i=-2}^{\infty} \sum_{j=0}^{j_{max}} \sum_{k=0}^{i+4} c_{ijk}^{(n)} \epsilon^i m_t^j \ln^k(m_t)$$

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High-Energy Expansion

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• Establish a system of differential equations for the master integrals in the variable m_t .

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Compare coefficients in ϵ and m_t to establish a linear system of equations for the $c_{ijk}^{(n)}$.



Calculation of the Master Integrals Deep Expansion



Compute an expansion around m_t = 0 by: Inserting an ansatz for the master integrals into the differential equation.

$$M_n(\epsilon, m_t \rightarrow 0) = \sum_{i=-2}^{\infty} \sum_{j=0}^{j_{\text{max}}} \sum_{k=0}^{i+4} c_{ijk}^{(n)} \epsilon^i m_t^j \ln^k(m_t)$$

• Compare coefficients in ϵ and m_t to establish a linear system of equations for the $c_{ijk}^{(n)}$.

Establish a system of differential equations for the master integrals in the variable m_t.

 Solve the linear system in terms of a small number of boundary constants using Kira and FireFly. [Klappert, Klein, Lange '19,'20]

Calculation of the Master Integrals



Compare coefficients in
$$\epsilon$$
 and m_t to establish a linear system of equations for the $c_{iik}^{(n)}$.

Solve the linear system in terms of a small number of boundary constants using Kira and FireFly. [Klappert, Klein, Lange '19.'20]

 $M_n(\epsilon, m_t
ightarrow 0) = \sum_{i=1}^{\infty} \sum_{j=1}^{j_{max}} \sum_{i=1}^{i+4} c_{ijk}^{(n)} \epsilon^i m_t^j \ln^k(m_t)$ i = -2 i = 0 k = 0

Master Integrals Results

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• Compute boundary values for $m_t \rightarrow 0$ and obtain an analytic expansion.

Calculation of the Master Integrals

\Rightarrow Why not utilize the differential equation in s or t?

High-Energy Expansion

Calculation of the Master Integrals

Deep Expansion

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- Establish a system of differential equations for the master integrals in the variable m_t .
- Compute an expansion around $m_t = 0$ by:
 - Inserting an ansatz for the master integrals into the differential equation.



Differential Equation in t



- We can always put one scale to unity, we choose $s \equiv 1$.
- We can use the differential equation in *t* in a similar manner.
- Boundary conditions are then only needed in the limit $m_t, |t| \rightarrow 0$.
- However, calculating the boundaries in the limit $m_t \rightarrow 0$ with full dependence on *t* turns out to be not harder than in the double limit m_t , $|t| \rightarrow 0$.

Differential Equation in t



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- Boundary conditions are then only needed in the limit m_t , $|t| \rightarrow 0$.
- However, calculating the boundaries in the limit $m_t \rightarrow 0$ with full dependence on *t* turns out to be not harder than in the double limit m_t , $|t| \rightarrow 0$.

\Rightarrow No benefit in utilizing the differential equation in *t*.

Calculation of the Master Integrals Boundary Conditions



How to obtain the boundary values?

• We start with the α representation of the diagram:

$$I_n = \int_0^\infty \left(\prod_{i=1}^n d\alpha_i \frac{\alpha_i^{\delta_i}}{\Gamma(1+\delta_i)}\right) \mathcal{U}^{-d/2} e^{-\mathcal{F}/\mathcal{U}},$$

with the Symanzik polynomials ${\mathcal U}$ and ${\mathcal F}.$

- We use expansion-by-regions [Beneke, Smirnov '98] and reveal the different regions with Asy.m [Pak, Smirnov '11].
- High-energy limit: $s, |t| \sim \chi^0, \, m_t^2 \sim \chi$
- In total we reveal 13 regions:
 - One hard region ($m_t = 0$), where master integrals are known [Smirnov, Veretin '00; Bern, Sixon, Smirnov '05].
 - 13 'soft' regions, where α parameters scale differently in χ .
- We calculate the expansion using Mellin-Barnes techniques.

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Calculation of the Master Integrals Boundary Conditions – Mellin-Barnes Techniques

• Symanzik polynomials: $(\alpha_{i_1...i_n} = \alpha_{i_1} + \cdots + \alpha_{i_n})$

 $\mathcal{U} = \alpha_{23}\alpha_{14} + \alpha_{1234}\alpha_5, \qquad \qquad \mathcal{F} = S\alpha_2\alpha_4\alpha_5 + T\alpha_1\alpha_3\alpha_5 + m_t^2\alpha_{12345}\mathcal{U}$

• 8 soft regions contribute for $m_t \rightarrow 0$: $(m_t^2 \rightarrow \chi m_t^2)$

 $\alpha_i \to \chi^{v_i^{(r)}} \alpha_i, \qquad \vec{v}^{(1)} = (0, 0, 0, 0, 1), \qquad \vec{v}^{(2)} = (0, 0, 1, 1, 0), \qquad \dots$

• After rescaling we can expand in χ , e.g.:

$$I_{5}^{(1)} = \int \left(\prod_{i=1}^{5} \frac{d\alpha_{i} \alpha_{i}^{\delta_{i}}}{\Gamma(1+\delta_{i})}\right) \mathcal{U}_{1}^{-d/2} e^{-\mathcal{F}_{1}/\mathcal{U}_{1}} \left[1 - \chi \left(m_{t}^{2} \alpha_{5} - S \frac{\alpha_{2} \alpha_{4} \alpha_{1234} (\alpha_{5})^{2}}{(\mathcal{U}_{1})^{2}} + \dots\right) + \dots\right]$$

with the expanded Symanzik polynomials

$$\mathcal{U}_1 = \alpha_{23}\alpha_{14}, \qquad \qquad \mathcal{F}_1 = \mathbf{S}\alpha_2\alpha_4\alpha_5 + \mathbf{T}\alpha_1\alpha_3\alpha_5 + \mathbf{m}_t^2\alpha_{1234}\mathcal{U}_1$$

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S = -s, T = -t

Boundary Conditions – Mellin-Barnes Techniques



Useful formula:

$$\int_{0}^{\infty} d\alpha \alpha^{a} e^{-A\alpha} = A^{-1-a} \Gamma(1+a),$$
$$\int_{0}^{\infty} d\alpha \alpha^{a} (A+B\alpha)^{b} = A^{1+a+b} B^{-1-a} \frac{\Gamma[1+a,-1-a-b]}{\Gamma(-b)},$$

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Boundary Conditions – Mellin-Barnes Techniques



Useful formula:

$$\int_{0}^{\infty} d\alpha \alpha^{a} e^{-A\alpha} = A^{-1-a} \Gamma(1+a),$$
$$\int_{0}^{\infty} d\alpha \alpha^{a} (A+B\alpha)^{b} = A^{1+a+b} B^{-1-a} \frac{\Gamma[1+a,-1-a-b]}{\Gamma(-b)},$$

$$\frac{1}{(A+B)^{\lambda}} = \int_{-i\infty}^{i\infty} \frac{dz}{2\pi i} \frac{B^z}{A^{\lambda+z}} \frac{\Gamma[-z,\lambda+z]}{\Gamma(\lambda)}, \quad \text{with} \quad \Gamma[x_1,x_2,\ldots,x_n] = \Gamma(x_1)\Gamma(x_2)\ldots\Gamma(x_n)$$



Boundary Conditions – Mellin-Barnes Techniques

• We can describe the expansion with one template integral:

$$T_{1,\{\delta_{1},\delta_{2},\delta_{3},\delta_{4},\delta_{5}\},\epsilon} = \int \left(\prod_{i=1}^{5} \frac{d\alpha_{i}\alpha_{i}^{\delta_{i}}}{\Gamma(1+\delta_{i})}\right) \mathcal{U}_{1}^{-d/2} e^{-\mathcal{F}_{1}/\mathcal{U}_{1}} = \frac{(m_{t}^{2})^{-\delta_{1234}-2\epsilon}}{S^{\delta_{5}+1}} \int \frac{dz_{1}}{2\pi i} \left(\frac{S}{T}\right)^{z_{1}} \frac{\Gamma[\delta_{23}+\epsilon,\delta_{14}+\epsilon,\delta_{2}-\delta_{5}-z_{1},-z_{1},\delta_{4}-\delta_{5}-z_{1},\delta_{1}+z_{1}+1,\delta_{3}\neq z_{1}+1,\delta_{5}+z_{1}+1]}{\Gamma[\delta_{1}+1,\delta_{2}+1,\delta_{3}+1,\delta_{4}+1,\delta_{5}+1,\delta_{23}-\delta_{5}+1,\delta_{14}-\delta_{5}+1]}$$

- We find up to 3-dimensional Mellin-Barnes integrals.
- The analytic continuation in δ_i and ϵ can be performed with MB.m [Czakon '05].
- The sum of all regions has to be free of poles in δ_i .
- \Rightarrow How to perform Mellin-Barnes integrals systematically?

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Mellin-Barnes Integrals Example

We find:

$$h_{3} = m_{t}^{-4\epsilon+2} \int \frac{dz_{1}}{2\pi i} \frac{\Gamma[-z_{1}, z_{1} - \epsilon + 2, -z_{1} + \epsilon - 1, z_{1} + 1, z_{1} + 1, z_{1} + \epsilon]}{\Gamma[2 - \epsilon, 2z_{1} + 2]}$$

• We use MB.m for the analytic continuation in ϵ :

$$I_3 = m_t^{-4\epsilon+2} e^{-2\epsilon\gamma_E} \left(-\frac{3}{2\epsilon^2} - \frac{9}{2\epsilon} - \frac{21}{2} - \frac{5\pi^2}{12} + I^{(MB)} + \mathcal{O}(\epsilon) \right)$$

• With the remaining integral:

$$I^{(MB)} = \int_{-1/7-i\infty}^{-1/7+i\infty} \frac{dz_1}{2\pi i} \frac{\Gamma[-z_1 - 1, -z_1, z_1, z_1 + 1, z_1 + 1, z_1 + 2]}{\Gamma(2z_1 + 2)}$$

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Mellin-Barnes Integrals Example

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$$\begin{split} \mathcal{I}^{(MB)} &= \int\limits_{-1/7 - i\infty}^{-1/7 + i\infty} \frac{dz_1}{2\pi i} \frac{\Gamma[-z_1 - 1, -z_1, z_1, z_1 + 1, z_1 + 1, z_1 + 2]}{\Gamma(2z_1 + 2)} \\ &= 4 + \frac{\pi^2}{6} + 2\sum_{k=0}^{\infty} \left(\frac{2k+1}{k}\right)^{-1} \frac{(4k^2 + 8k + 3)[S_1(k) - S_1(2k)] - 4(k+1)}{(2k+1)(2k+2)(2k+3)^2} \end{split}$$



Summation over residue sum can be done analytically with HarmonicSums [Ablinger et al. '10-], Sigma and EvaluateMultiSums [Schneider '07-].

The (inverse) binomial sums we encounter sum to special constants, e.g.:

$$\sum_{k=0}^{\infty} \xi^k \binom{2k+1}{k}^{-1} \frac{1}{3+2k} = \frac{2}{\xi\sqrt{(4-\xi)\xi}} \int_0^{\xi} dt \sqrt{(4-t)t} - 1 \stackrel{\xi \to 1}{=} \frac{4\pi}{3\sqrt{3}} - 2$$



Mellin-Barnes Integrals Example

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- Most complicated boundary condition: $G_4(1,1,1,1,1,1,1,-1,-1)$
- The irreducible numerators can be handled by starting from the topology with all 9 lines.
- We end up with a large number of Mellin-Barnes integrals: one-dimensional | two-dimensional | three-dimensional

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Taking residues and summation can be automatized.







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Boundary Conditions – Pitfalls





not in agreement with numerical evaluation.

• Problem: integral does not fall off fast enough for $|z_2| \rightarrow \infty$.

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Boundary Conditions – Pitfalls



- Problem: integral does not fall off fast enough for $|z_2| \to \infty$.
- We can solve this problem with regularization:

$$\begin{split} & U = \int_{-1/7 - i\infty}^{-1/7 + i\infty} dz_2 \, \xi^{z_2} \, \frac{z_2^8 \Gamma^2 (-z_2) \Gamma^2 (z_2)}{(z_2 + 1)^3 (z_2 + 2)^3} = -\sum_{k=0}^{\infty} \xi^k \left(\frac{3k^5 (4 + 3k)}{(1 + k)^4 (2 + k)^4} + \frac{k^6}{(1 + k)^3 (2 + k)^3} \ln(\xi) \right) \\ & = \sum_{k=0}^{\infty} \xi^k \left(\frac{3k^5 (4 + 3k)}{(1 + k)^4 (2 + k)^4} + \left[1 - \frac{(2 + 3k)(4 + 12k + 15k^2 + 9k^3 + 3k^4)}{(1 + k)^3 (2 + k)^3} \right] \ln(\xi) \right) \\ & \stackrel{\xi \to 1}{=} -18\zeta_3 - \frac{3\pi^2}{3} - \frac{21\pi^4}{10} + 240 + 1 \end{split}$$

Alternative approach: high precision numerical evaluation in combination with PSLQ [Ferguson, Bailey '92].

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Master Integrals Results

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Master Integrals Results



- We obtain analytic expressions of all 140 master integrals up to $\mathcal{O}(m_t^{120})$.
- The final result can be expressed via harmonic polylogarithms [Remiddi, Vermaseren '99]

 $H_{0}(-t/s), H_{1}(-t/s), H_{0,1}(-t/s), H_{0,0,1}(-t/s), H_{0,1,1}(-t/s), H_{0,0,0,1}(-t/s), H_{0,0,1,1}(-t/s), H_{0,1,1,1}(-t/s), H_{0,1,1,1}($

and transcendental numbers

$$\pi, \ln(3), \sqrt{3}, \zeta_2, \zeta_3, \psi^{(1)}(1/3), \ln\left[\text{Li}_3(i/\sqrt{3})\right]$$

• We also extended the calculation of the master integrals with massless internal line up to $\mathcal{O}(m_t^{120})$.

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Padé-Improved High-Energy Expansion



The master integrals for both methods are computed as an expansion in $m_t \ll s$, |t|.

The expansions diverge for \sqrt{s} \sim 750GeV ("A"), \sqrt{s} \sim 1000GeV ("B").

The situation can be improved using Padé Approximants:

Approximate a function using a rational polynomial

$$f(x) \approx \frac{a_0 + a_1 x + a_2 x^2 + \dots + a_n x^n}{1 + b_1 x + b_2 x^2 + \dots + b_m x^m}$$

where a_i , b_j coefficients are fixed by the series coefficients of f(x).

We compute a set of various Padé Approximants:

- combine to give a central value and error estimates
- a deeper input expansion \Rightarrow larger $n + m \Rightarrow$ smaller error
- here, m_t¹²⁰ expansion allows for very high-order Padé Approximants

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Master Integrals Results Padé Improvement









$$p_T^2 = \frac{tu - m_h^4}{s}$$

- Lower order Padé approximantions cannot reach low values of p_T.
- For QCD corrections expansions up to m_t^{32} were available: $p_T \gtrsim 150 \, {\rm GeV}$
- With expansions up to m_t^{120} we reach: $p_T \gtrsim 120$ GeV.
- Error estimate from Padé approximations is reliable.





$$p_T^2 = \frac{tu - m_h^4}{s}$$

- Lower order Padé approximantions cannot reach low values of p_T.
- For QCD corrections expansions up to m_t^{32} were available: $p_T \gtrsim 150 \, {\rm GeV}$
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Comparison to the $m_H \rightarrow 0$ Expansion





Approach A:

- middle line massless $m_H^{\rm int} \approx 0$
- calculated in the context of QCD corrections

[Davies, Mishima, Steinhauser, Wellmann '18, '19]

Approach B:

• middle line massive $m_H^{\rm int} \approx m_t$







0.3 0.6 0.2 0.5 0.1 0.4 real part, approach A real part, approach B 0.0 0.3 0.2 -0.1-0.20.1 -0.3 0.0 -0.4 -0.1-0.5-0.2 -0.61 -0.3350 400 450 500 550 600 350 400 450 5**0**0 550 600 $\sqrt{s}(GeV)$ $\sqrt{s}(GeV)$ Approach B: threshold at $\sqrt{s} = 3m_t = 519 \,\text{GeV}$

Comparison with Approach A

Approach A: threshold at $\sqrt{s} = 2m_t = 346 \,\text{GeV}$

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Form Factor Results



Comparison with Approach A



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Master Integrals Results

Form Factor Results

Conclusion and Outlook

Renormalization

The form factors require UV renormalization (they are IR finite):

MS renormalization of the top quark mass,

$$m_t^0 o \overline{m}_t \left[1 + rac{\alpha_t}{\pi} rac{1}{\epsilon} \left(rac{3}{16} + rac{N_C}{2} rac{\overline{m}_t^2}{m_H^2}
ight)
ight]$$

 \blacksquare LO has no δ expansion, so NLO δ terms must already be finite \checkmark

The second term in (\cdots) renormalizes the tadpole diagrams,

computed, but not included in the following plots.



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Universität High-Energy Expansion and Padé Approximation Zürich[™] Re[F1] 4 mhs2 mts15 delta3 mhs2 mts16 delta3 mhs2 mts56 delta3 mhs2 mts57 delta3 Pade (mts16) ····· Pade (mts57) $\operatorname{Re}(F_{box1})$, fixed $\cos \theta = 0$, <u>4</u>00 √s 600 800 1000 expansion "B" 1200 1400 $(to (m_H^2)^2 \delta^3 (m_t^2)^{\{15,16,56,57\}}):$ m_t expansion diverges (strongly) around \sqrt{s} \sim 1000GeV Introduction High-Energy Expansion Calculation of the Master Integrals Master Integrals Results Form Factor Results Conclusion and Outlook 000000



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Convergence of Asymptotic Expansion ("A")

Kay Schönwald: Yuakawa corrections to $qq \rightarrow HH$ 39/43 18.10.22



Comparison of "A", "B" Expansions





Form Factors at Fixed p_T





Conclusion and Outlook

Expansions "A" and "B" agree for p_T values as small as 120 GeV.

 deep expansions of the MIs required, for small Padé errors

Form Factor Results

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Conclusion

Conclusions:

First step towards electroweak corrections to double Higgs production:

- more difficult than the QCD contribution (extra internal scale)
- expansion allows us to compute them

High-energy expansion:

- Padé-based approximation to improve expansion
- good description of (partial) form factors for $p_T\gtrsim$ 120GeV
- two different expansion methods, which give equivalent results
- \blacksquare deeper exp. of MIs compared to QCD papers \rightarrow better Padé



43/43 18.10.22 Kay Schönwald: Yuakawa corrections to $gg \rightarrow HH$

Conclusion

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Outlook:

Introduction

- Apply calculation strategy to the full electroweak corrections.
 - \Rightarrow This will include also non-planar sectors.

High-Energy Expansion

Explore complementary expansions to cover the whole kinematic range.

Calculation of the Master Integrals

Master Integrals Results



