

Dark matter search with the XENON100 experiment

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Joint EP/PP Seminar, CERN, 19.04.2010



Outline

- 1 Introduction
- 2 The XENON experiment
- 3 Results from calibration sources
- 4 First XENON100 results
- 5 R&D at UZH
- 6 Summary

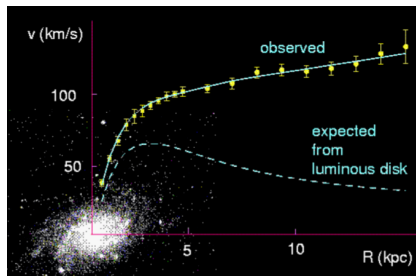
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Indications from astronomy

Star rotation curves

- Measurement: 21 cm H-line
- Dark matter halo explanation

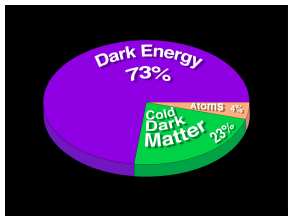


Bullet cluster

- Indirect dark matter evidence
- Baryonic matter in red (X-rays)
- Matter distribution in blue (gravitational lensing)

+ large scale structures, WMAP data...

Dark matter theoretical predictions



DM particle candidates:

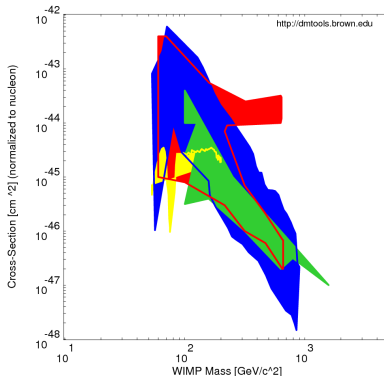
- Sterile neutrinos
- Axions
- Kaluza-Klein states
- SUSY particles
 - Neutralino χ
- Little Higgs ...

WIMP

(Weakly Interacting Massive Particle)

“Enhancement” of CMSSM predictions around $10^{-45 \pm 2} \text{ cm}^2$

Yellow: A. Pierce, Finely Tuned MSSM
Red: Baer et. al 2003
Green: rSUGRA, $M_3=M_1$, $0.6 < \tan\beta < 1$, $5 < \tan\beta < 50$
Blue: Ellis et. al 2005 CMSSM ($\mu > 0$, pion Sigma=64 MeV)



WIMP search

- Indirect detection



$$\chi\chi \rightarrow e^+e^-, p\bar{p}$$

- Direct detection



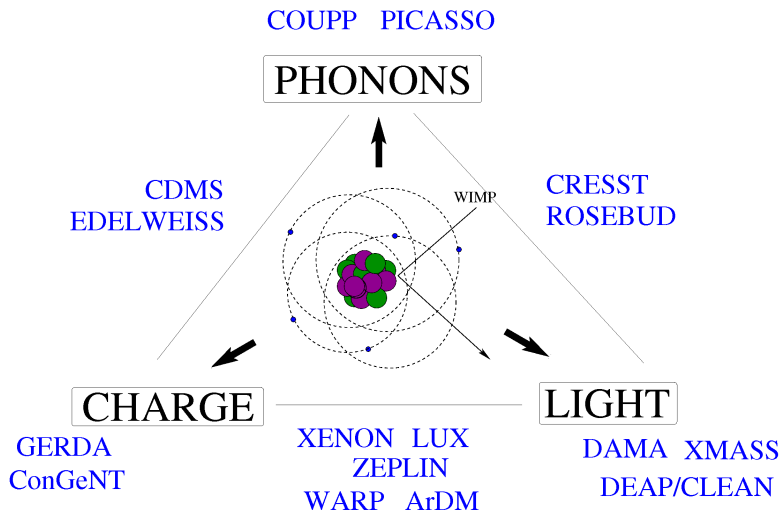
$$\chi N \rightarrow \chi N$$

- Production at LHC

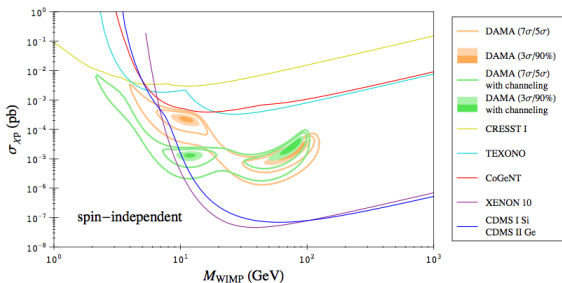
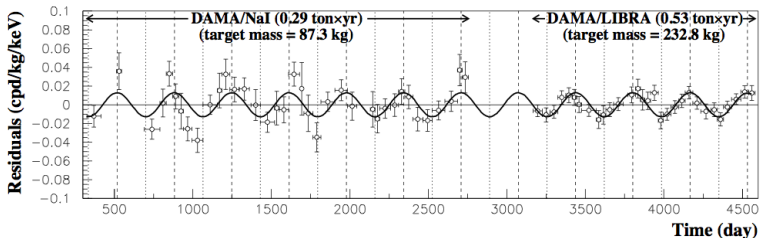


$$p + p \rightarrow \chi + \text{a lot}$$

Direct detection experiments



DAMA compared to other experiments



- Significant seasonal oscillation in the 2-6 keV energy bin
- Tension with other DM experiments

→ Plot from C.Savage, G.Gelmini, P.Gondolo and K. Freese (2009)

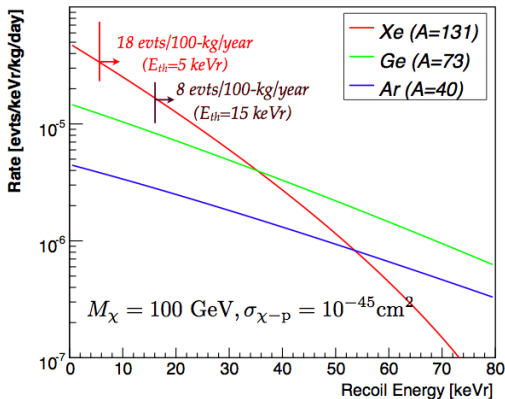
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Xe

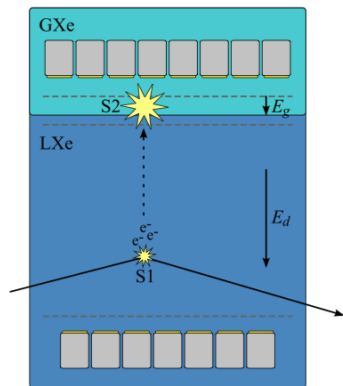
Detection via scatter off nuclei

WIMP Scattering Rates



- Self-shielding
→ High stopping power
- 178 nm UV photons
→ No wavelength-shifter
- Simple cryogenics
~ 180 K = - 93°
- High atomic mass $Z \sim 131$
→ spin-indep. interactions
- ^{129}Xe and ^{131}Xe
→ spin-dep. interactions

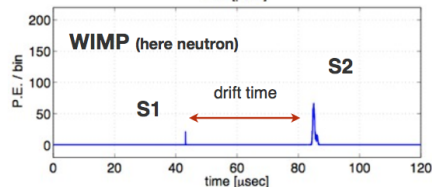
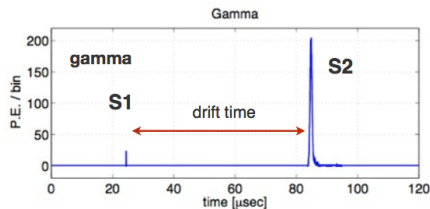
Two phase noble gas TPC



Electron recombination is stronger for nuclear recoils

→ Electron- / nuclear recoil discrimination

- Scintillation signal (S1)
- Charges drift to the liquid-gas surface
- Proportional signal (S2)



XENON experiment



- Laboratori Nazionali del Gran Sasso (Italy)
- 3 500 m.w.e. shielding

- **XENON10**: 15 kg active volume
 - Finished: No evidence for DM

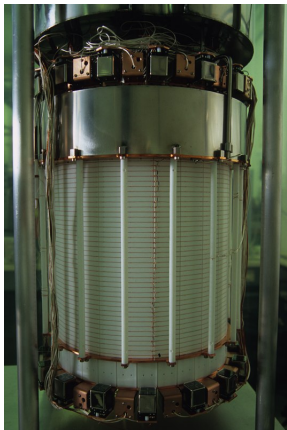
J. Angle *et al.*, Phys. Rev. Lett. 100, 021303 (2008)
J. Angle *et al.*, Phys. Rev. Lett. 101, 091301 (2008)
J. Angle *et al.*, Phys. Rev. D80, 115005 (2009)

- **XENON100**: 65 kg active volume
 - Currently taking science data



XENON100 detector

- 30 cm drift length and 30 cm $\varnothing$
- 165 kg total (30-50 kg fiducial volume)
- $\sim 100\times$ less background than XENON10
- Improved shielding
- Material screening and selection
- Cooling (PTR) outside the shield
- Active liquid xenon veto



1 inch PMTs

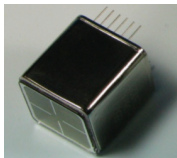


30 cm $\varnothing$ meshes

XENON100 detector



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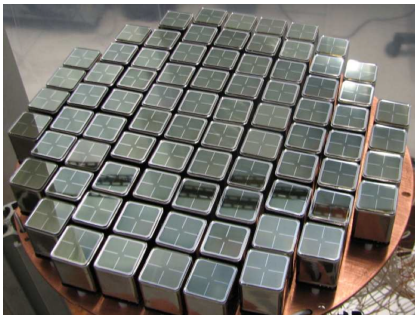


1 inch PMTs



30 cm $\varnothing$ meshes

Light and charge read out

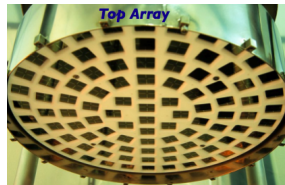


gamma event localized



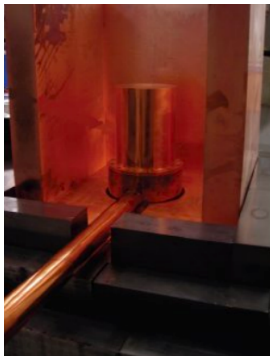
Top PMT array

- High quantum efficiency PMTs in the bottom array (>32% @178 nm)
- 3 Dim. position reconstruction
 - XY from light pattern in the PMTs
 - Z from the drift time
- 2 mm resolution in XY and in Z

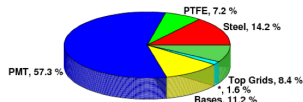


Background prediction

Material screening underground with
a 2.2 kg HP Ge detector



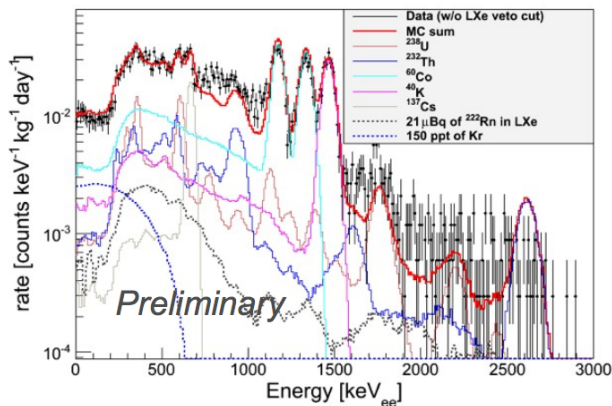
- Gamma background expected:
 - $5 \cdot 10^{-3}$ evts/kg/keV/d
(before discrimination cuts)



- Neutron bg from simulations:
 - < 1 event/year
 - 2/3 from radioactivity and
1/3 muon-induced

- Removal of ^{85}Kr : distillation column
 - Kr/Xe $\sim$ ppm-ppb commercially available
 - Measurement in XENON100 after purification:
 - currently ~ 140 ppt via delayed gamma-beta coincidence

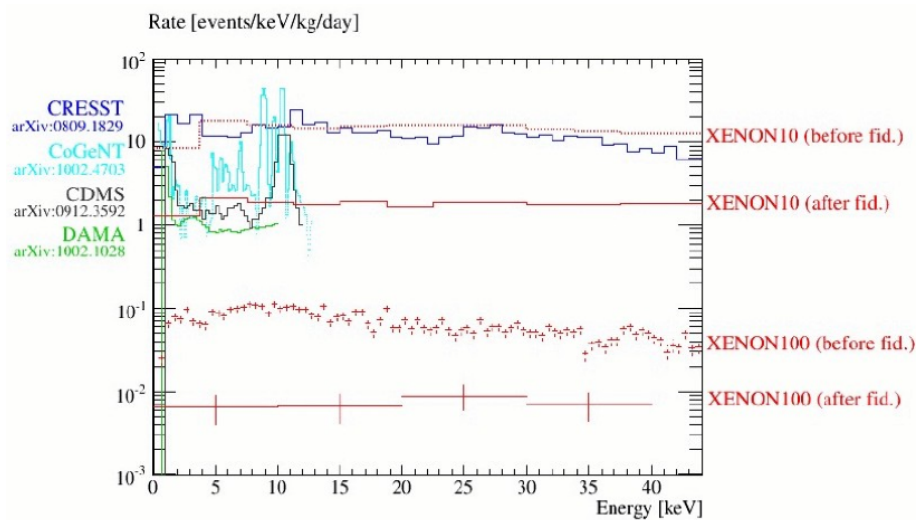
Measured background spectrum



- Background at the level of predictions
- No tuning of the Monte Carlo
- The measured single scatter rate below 100 keV_{ee} is 10^{-2} evts/kg/keV/d without veto cut
- **Factor 100 less** than in XENON10 achieved!!

→ currently optimizing the data/MC comparison

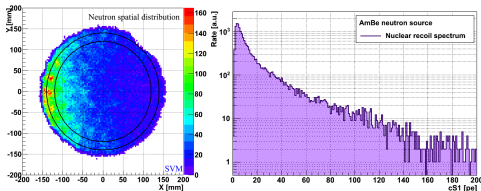
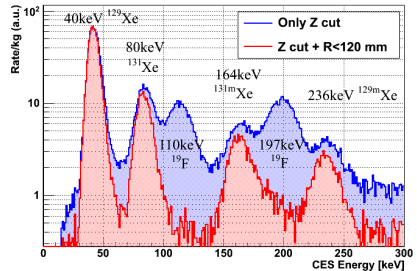
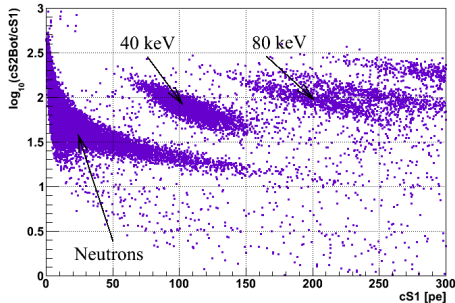
Lowest ever measured background rate



Outline

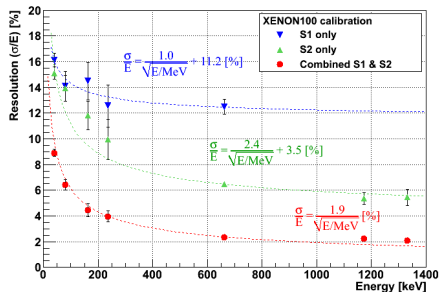
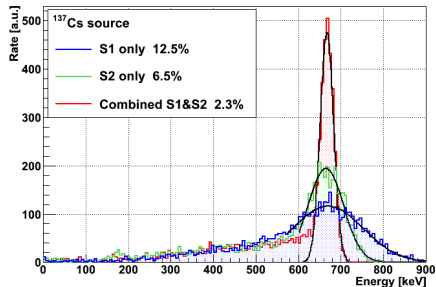
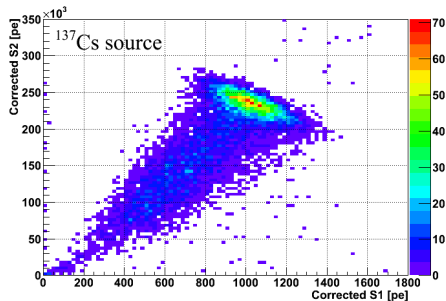
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Neutron calibration



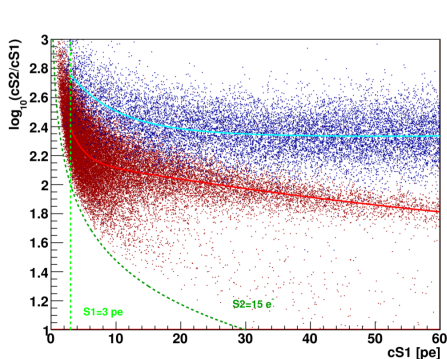
- Source: AmBe with 220n/s
- Determination of nuclear recoil band
- Further lines from inelastic recoils in xenon

Calibration with gamma sources

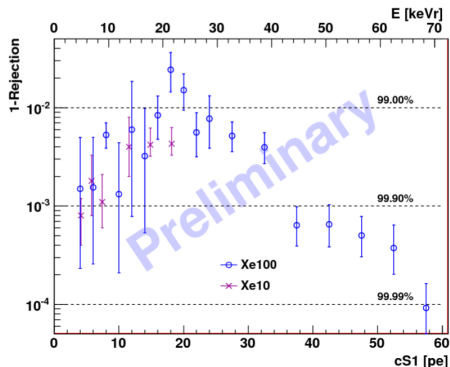


- Energy dependence of resolution in S1, S2 and CES
- CES: combined energy scale
- anticorrelation between S1 and S2

PRELIMINARY discrimination



- Band determination using:
 ^{60}Co and AmBe data

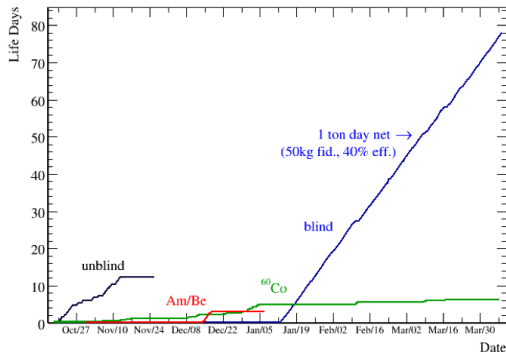


- Discrimination similar to
XENON10 @ about 50% NR
acceptance

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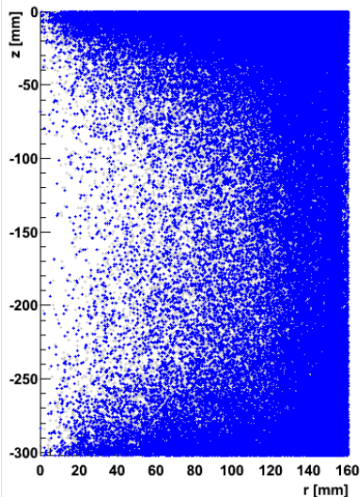
Overview of the data taking:



- Analysis of **unblinded** data
- Main data sample (blinded) not yet analyzed!

- 11.17 life days
- Data selection based on **stable conditions**:
 - no activation
 - no HV problems
 - no radon
- Period:
October-November 2009
- Cuts defined on calibration data: **AmBe** and ^{60}Co

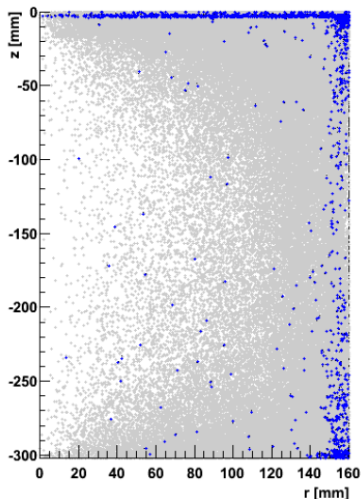
Basic cuts



→ Spatial distribution of the events after:

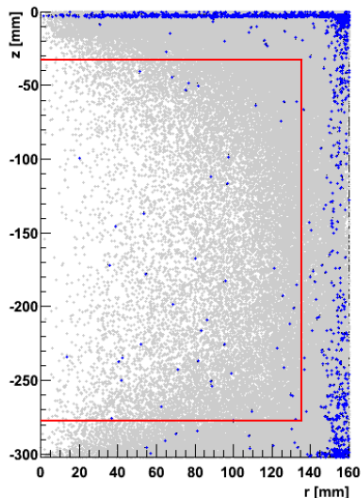
- Signal/noise ratio cut
- Single S1 peak: remove accidental coincidences
- Single S2 peak: select single scatter events
- Remove events in gas phase
- Apply active veto cut

Energy cut



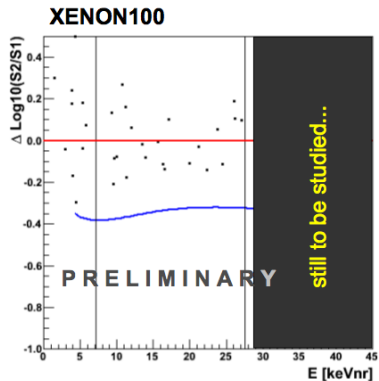
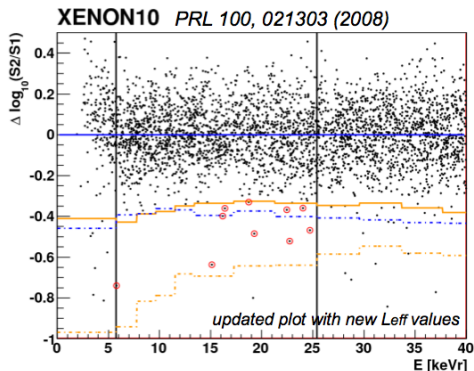
- **Energy cut:** Select events with energies lower than 28 keVnr (keV nuclear recoil equivalent)
→ defined region of interest for elastic scattering of WIMPs in xenon
- **Self-shielding:** Most of the low energy events are located close to the edges of the detector

Fiducial volume



- 3-D position reconstruction:
allows for a selection of the
inner part of the detector
- Current fiducial volume:
cylindrical shape
with 40 kg mass
- will be further optimized

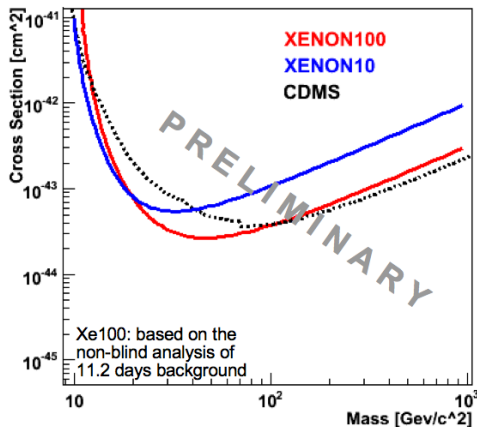
Particle discrimination



- 'Background free': in the 11.17 days after discrimination
 - Comparison to XENON10: for approximately the same exposure
- much cleaner detector

Limit from unblinded data

- **Spin independent limit:** for standard halo parameters
 - using the energy region $\sim (7 - 28) \text{ keVnr}$
- **Excellent sensitivity:** even for few days of data
 - Results will be published soon
- Much more data recorded in blind mode
- + analysis in the high nuclear-recoil energy region



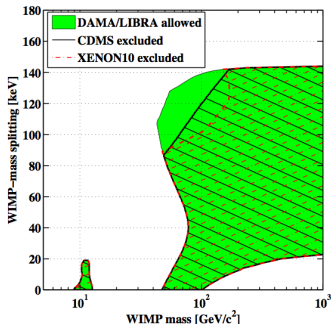
Interpretation of DAMA signal as iDM

- Inelastic dark matter model (D. Tucker-Smith and N. Weiner)

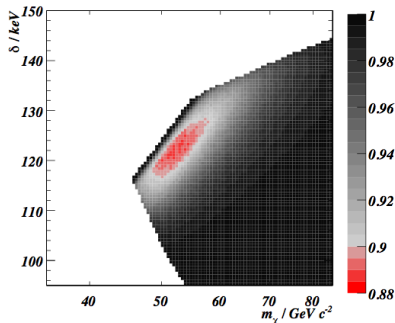
- WIMP scatter to an excited state

- Minimum relative speed: $v_{min} = \frac{1}{\sqrt{2m_N E_R}} \left(\frac{m_N E_R}{\mu_N} + \delta \right)$

→ Expected WIMP rates for Ge/Xe at higher nuclear recoil energies



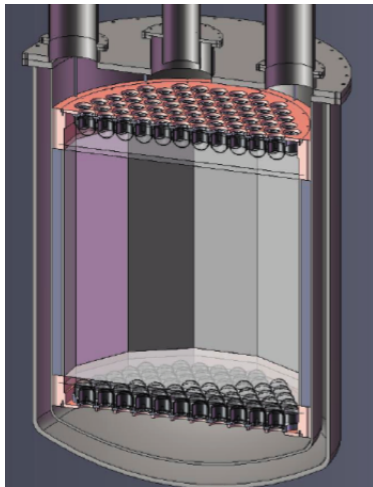
CDMS detector: 230 g of germanium
arXiv: 0912.3592 [astr-ph]



Zeplin: 6.5 kg liquid xenon
Latest result arXiv:1003.5626v2 [hep-ex]

- XENON100 results can check the remaining allowed region

Future: XENON1t

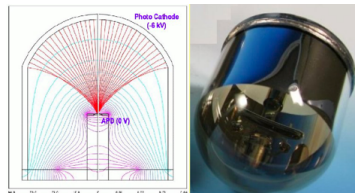
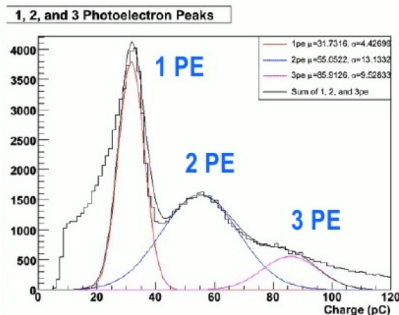


- 1 ton fiducial mass
(total of 2.2 ton LXe)
 - Drift length = ~ 90 cm
 - 100x background reduction
 - Muon veto
 - Copper/titanium cryostat
 - New photo-detectors?
- Schedule: 2010 - 2015 ?
- New collaborators
- Currently working on MC simulations and design +secure funding

Photosensors

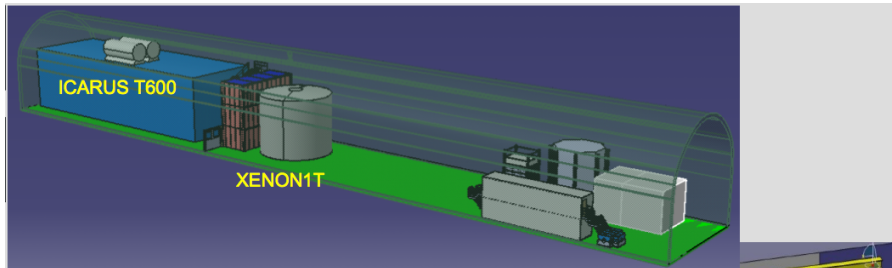
- **QUPIDS** for light readout
(**QU**artz **P**hoton **I**ntensifying **D**etector)

- Ultra-low radioactivity
(~ 0.1 mBq)
- High QE and high SPE resolution
- Development UCLA & Hamamatsu



- First test at UCLA
 - QUPID working in liquid xenon!
- single electron response

Location under discussion ...

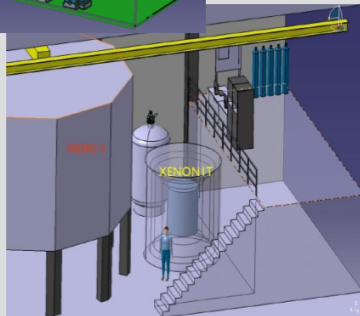


XENON1T @ LNGS (Hall B)

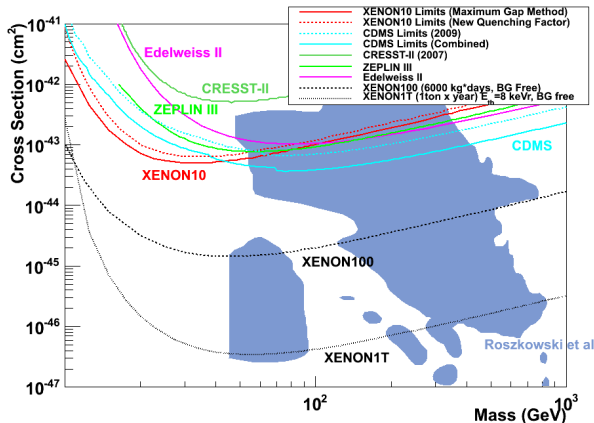
→ 4 m water shield

XENON1T @ LSM

→ solid shield (55cm poly,
20cm Pb, 15cm poly,
2cm ancient Pb,
>99% muon veto)



XENON sensitivity



- XENON100 sensitivity for 6 000 kg days (200 d × 30 kg bg free)
- Capability to detect about 10 events for 100 GeV mass for a WIMP-nucleon cross section of $\sim 10^{-44} \text{ cm}^2$ within 2010

Outline

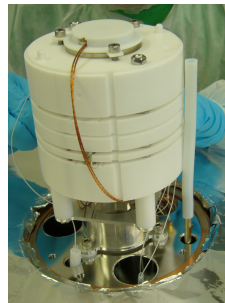
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R&D with the Xürich detector

Two-phase xenon detector at UZH



- Investigate light and charge production by different types of particles (electron/neutron)
- Test of calibration sources for xenon detectors

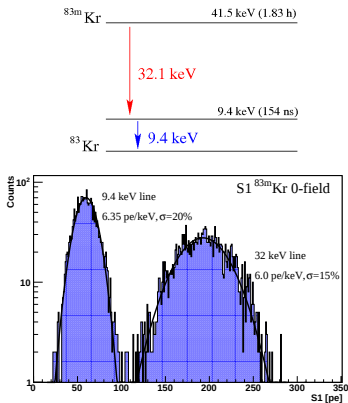


- Target mass: ~ 0.1 kg Xe
- Volume: 3 cm drift length and 3.5 cm diameter
- Two R9869 PMTs
- 10 pe/keV in single phase
- 6 pe/keV in double phase

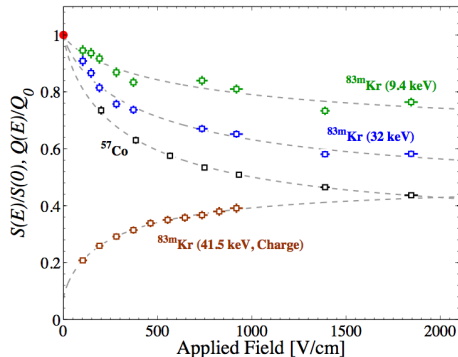
Low energy calibration of xenon detectors

- ^{83m}Kr calibration source:

- EC decay-product of ^{83}Rb
- Lines at 9.4 and 32.1 keV
- Uniform distribution



- ^{83m}Kr calibration planned in XENON100



A. Manalaysay *et al*, under review at Rev. Sci. Instrum. (2010)

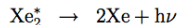
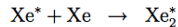
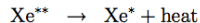
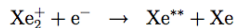
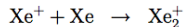
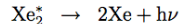
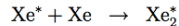
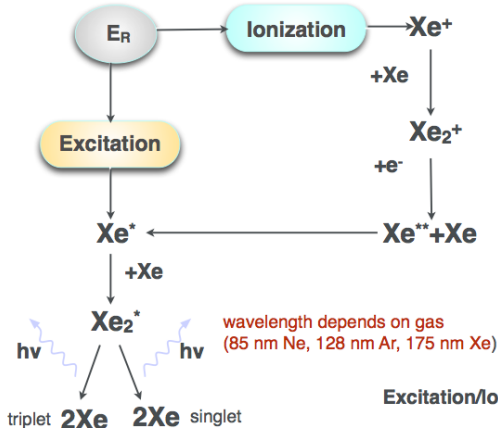
- Future: n-generator facility

- monoenergetic D-D fusion source
- charge and light yield of NR

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- **Liquid xenon** is a promising detector material to discover dark matter
 - Large nucleus (A^2 enhancement on σ)
 - Dual-phase: particle discrimination
 - Self-shielding (large detectors)
- **XENON100** is taking dark matter data
 - Design background level achieved:
 - lowest ever measured!
 - First unblinded data analyzed
 - new limit to be published!
- **R&D** ongoing to test new calibration sources
- **XENON1t** currently under design



Excitation/Ionization depends on dE/dx !

$\Rightarrow$ discrimination of signal (**WIMPs** $\Rightarrow$ NR)
and (most of the) background (**gammas** $\Rightarrow$ ER)!

time constants depend on gas
(few ns/15.4 μ s Ne, 10ns/1.5 μ s Ar, 3/27 ns Xe)